

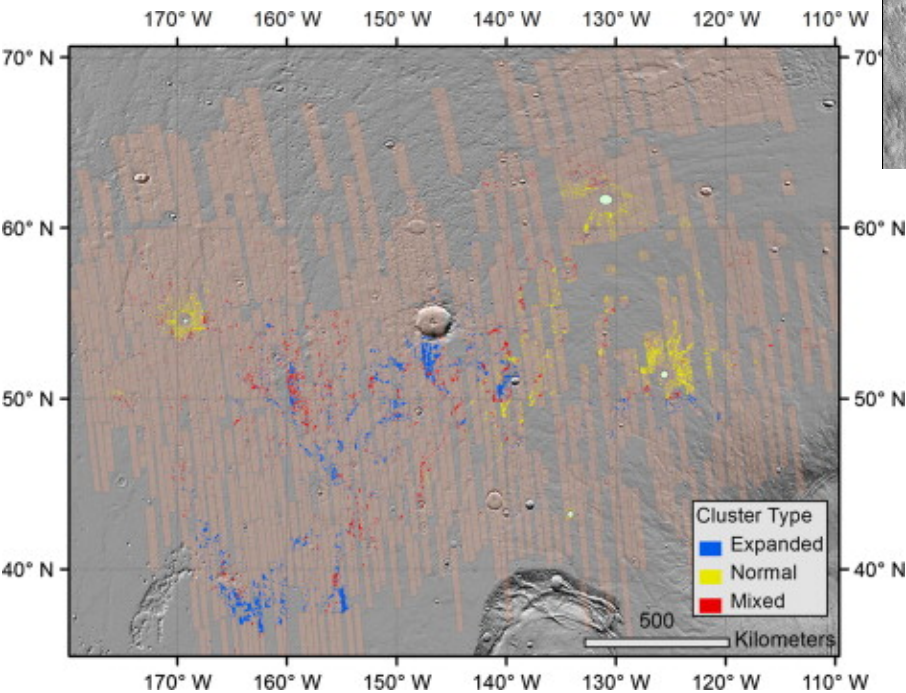
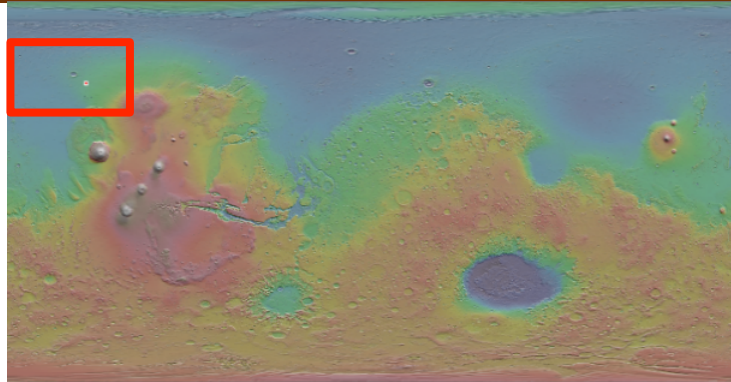
Arcadia Planitia: Acheron Fossae and Erebus Montes Workshop Abstract #1011

Donna Viola¹ (dviola@lpl.arizona.edu), Alfred S.
McEwen¹, & Colin M. Dundas²

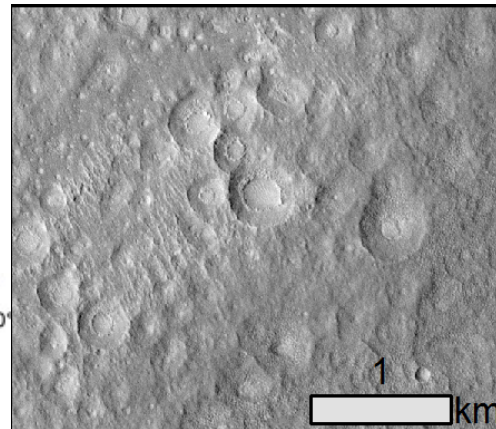
¹University of Arizona, Department of Planetary
Sciences, ²USGS Astrogeology Science Center

Motivation: Shallow Subsurface Water Ice

1st EZ Workshop for Human Missions to Mars



Expanded Secondary Craters



These craters require substantial **excess ice** in order to form. They suggest that the Arcadia Planitia ice sheet is tens of millions of years old.

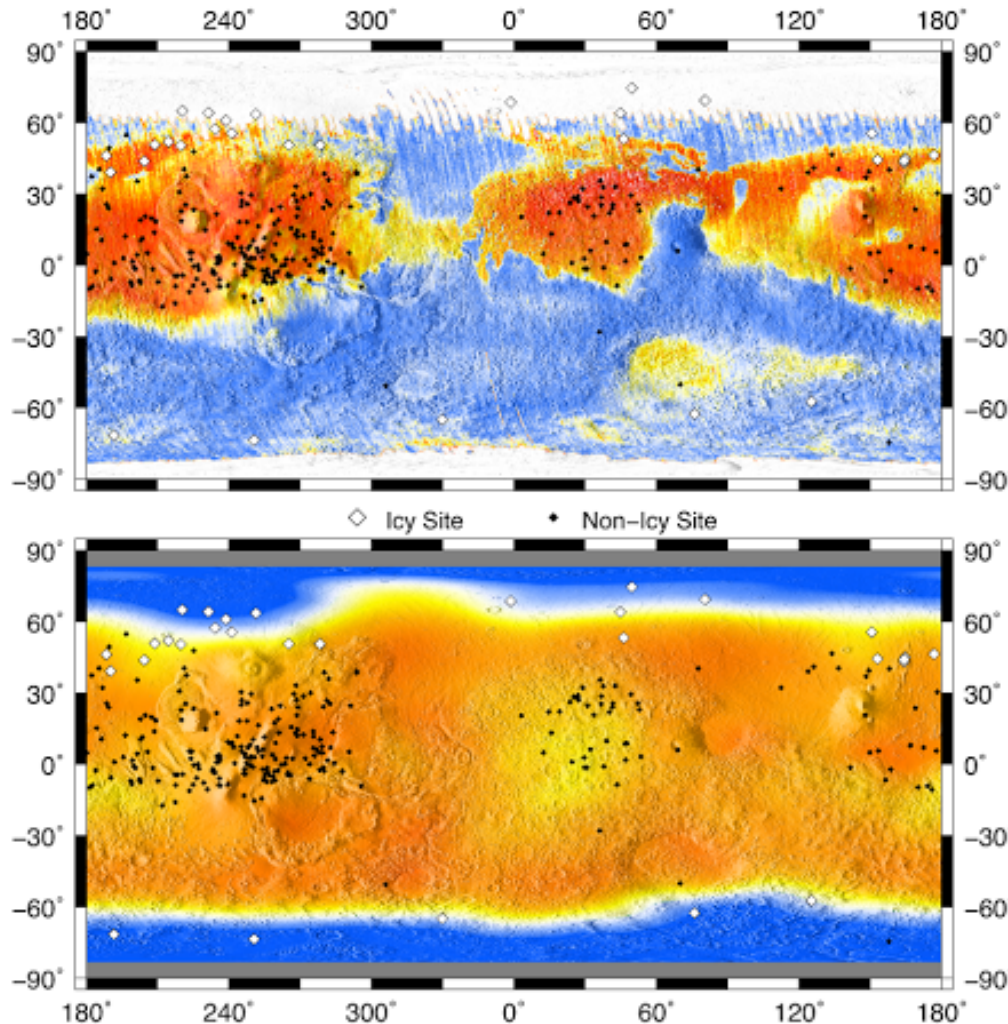
Viola, et al. (2015). *Icarus* 248: 190-204.

Arcadia Planitia is one of the few regions where abundant shallow ice is present at relatively low latitude.

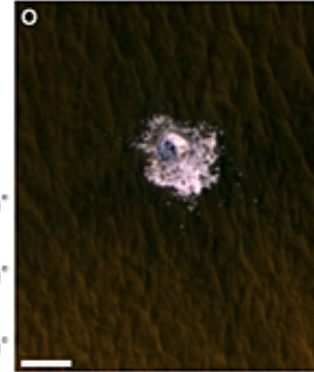
Motivation: Shallow Subsurface Water Ice

1st EZ Workshop for Human Missions to Mars

Map of recent ice-exposing impact craters



Dust cover index basemap (**red** = dusty, **blue** = less dusty)



These craters expose “**excess**” ice (Dundas et al., 2015)– which is almost entirely free of dust (>99% water ice).

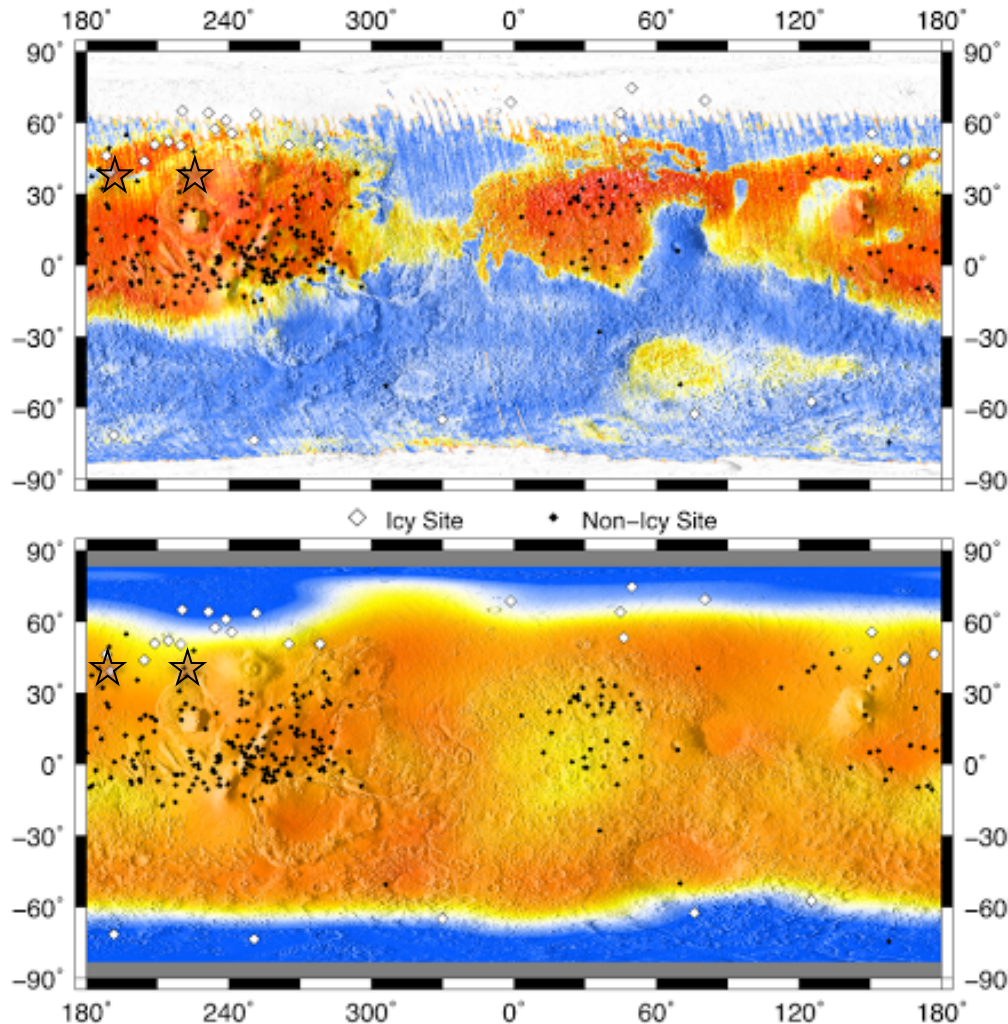
Water-equivalent hydrogen basemap (**blue** = more ice, **orange** = little/no ice)

Arcadia Planitia is one of the few regions where abundant shallow ice is present at relatively low latitude.

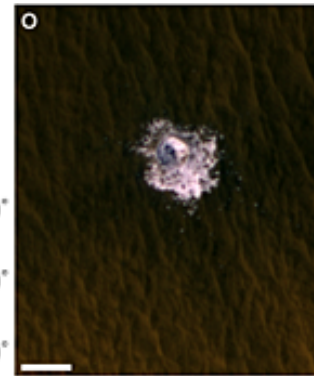
Motivation: Shallow Subsurface Water Ice

1st EZ Workshop for Human Missions to Mars

Map of recent ice-exposing impact craters



Dust cover index basemap (**red** = dusty, **blue** = less dusty)



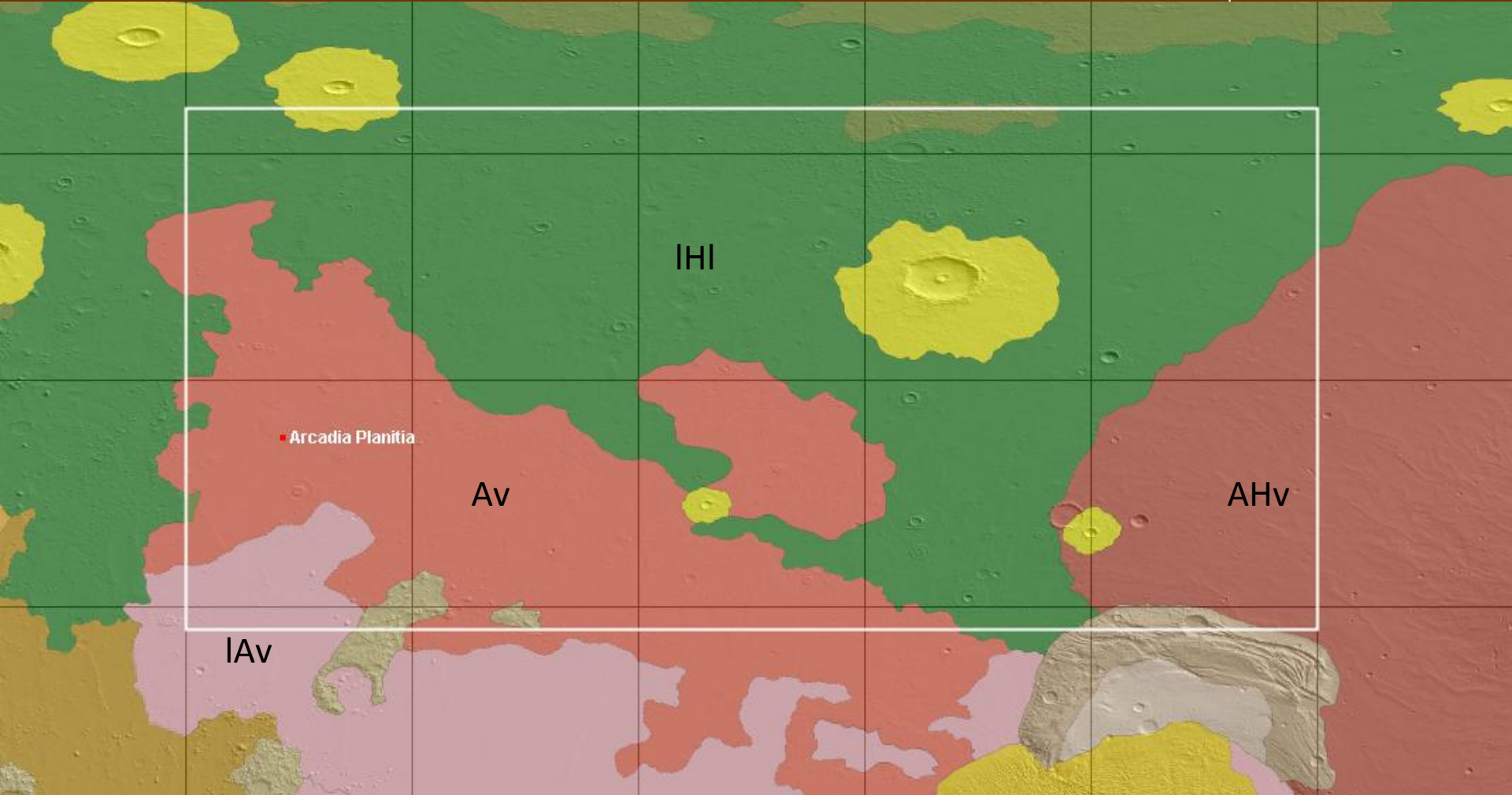
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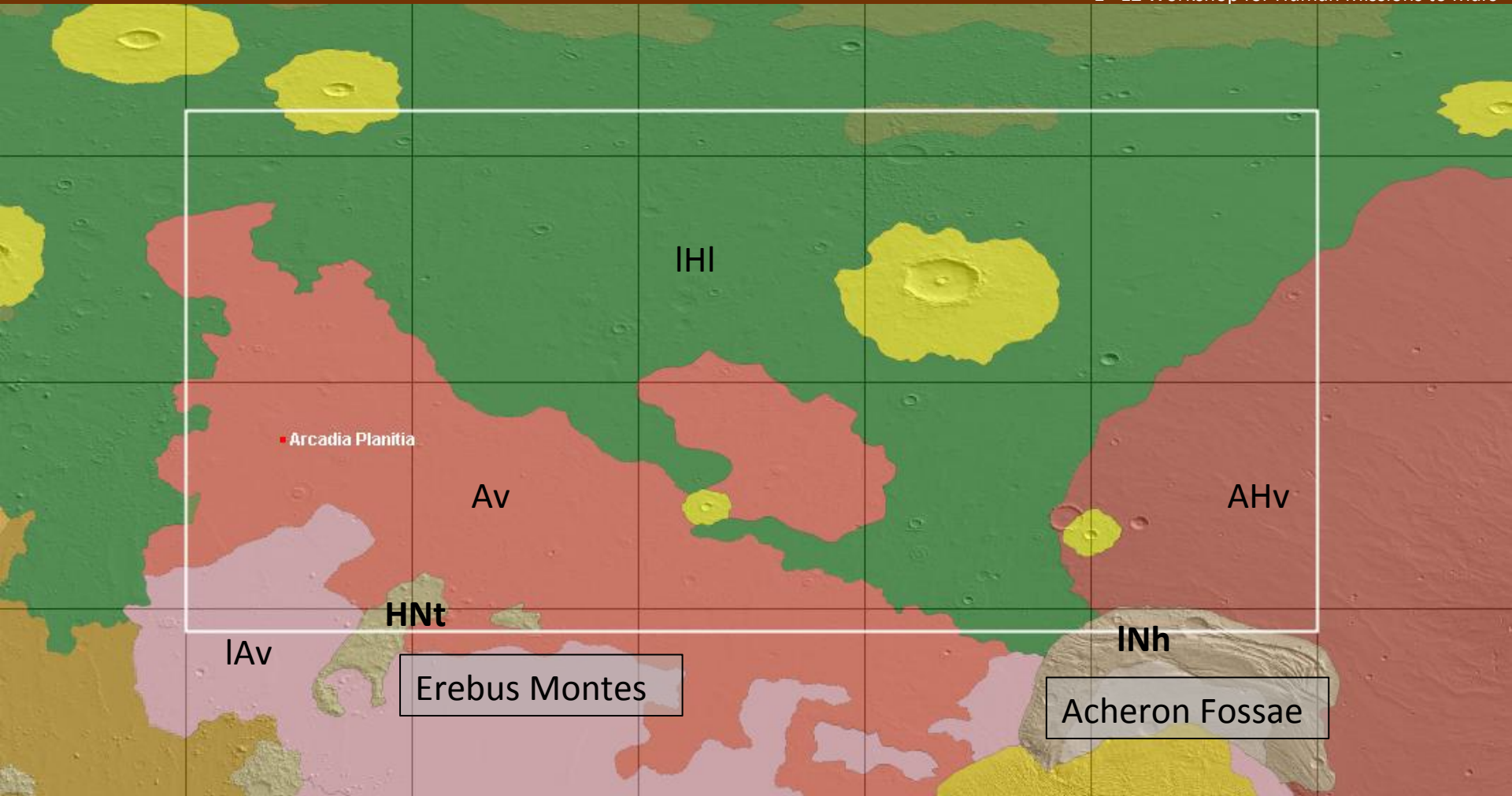
Arcadia Planitia: Geologic Context

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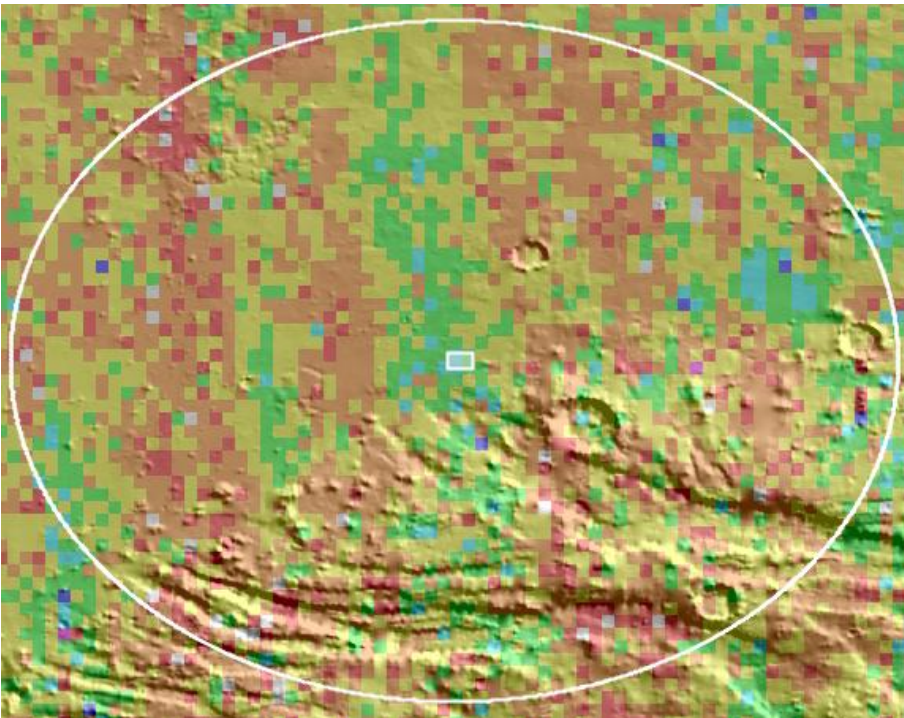
Arcadia Planitia: Geologic Context

1st EZ Workshop for Human Missions to Mars

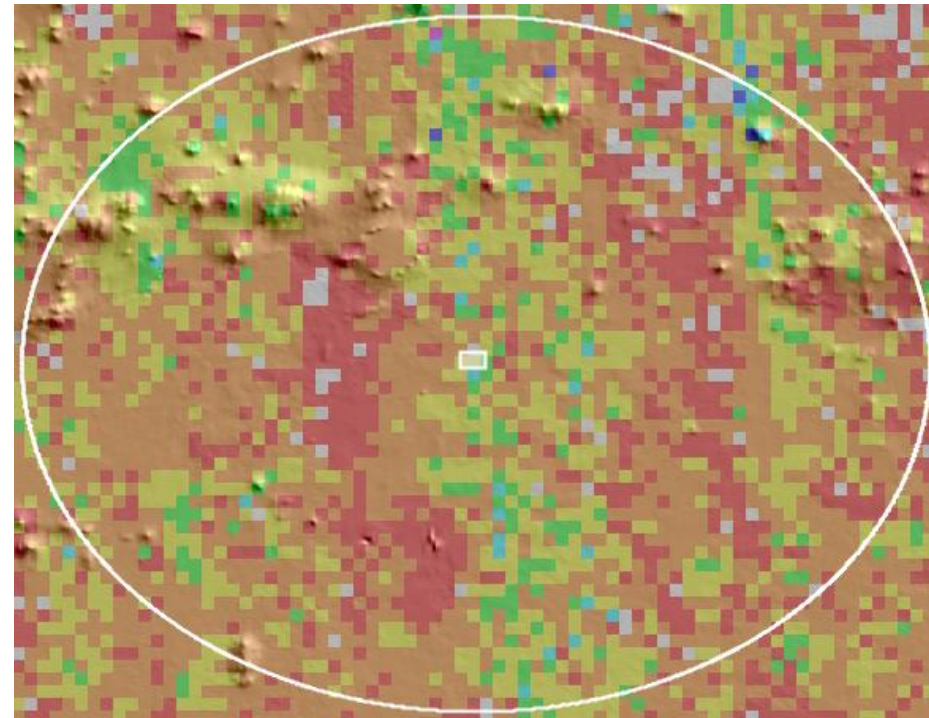


Arcadia Planitia: Challenges

1st EZ Workshop for Human Missions to Mars



Acheron Fossae

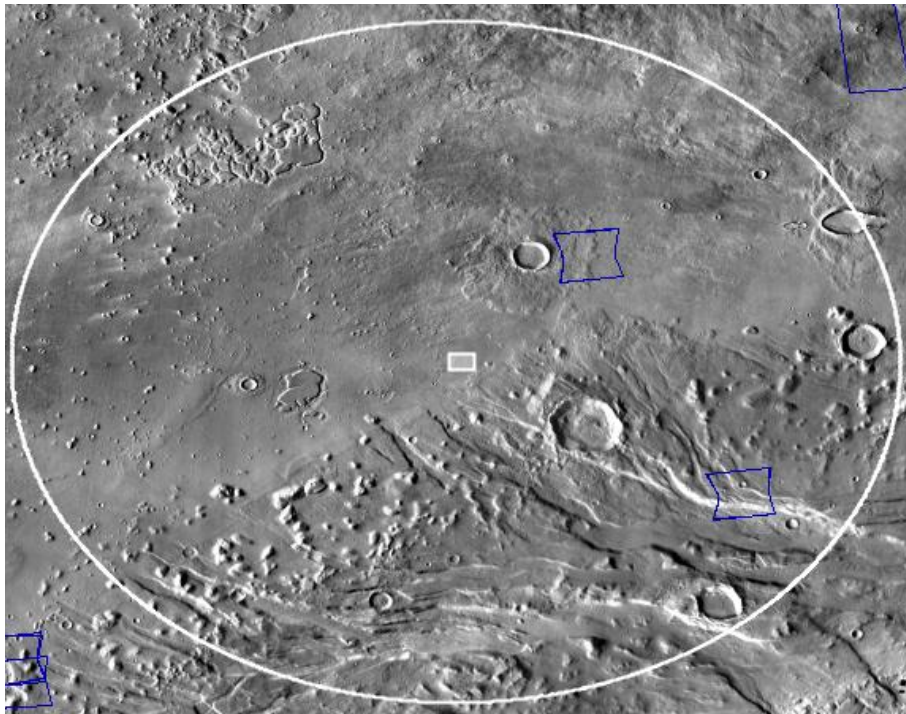


Erebus Montes

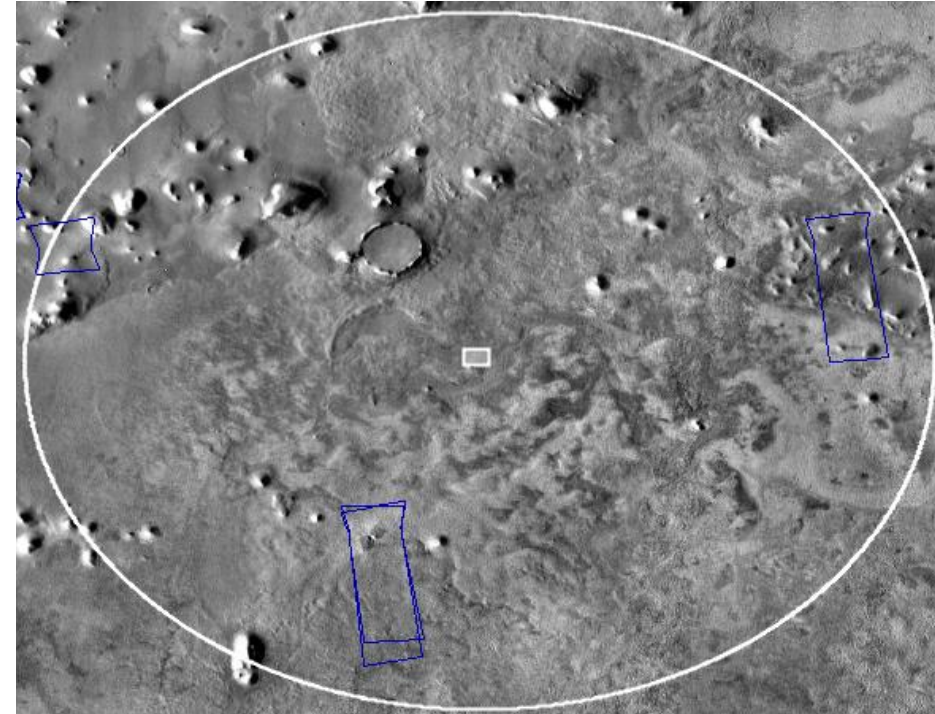
Dust: Arcadia Planitia is fairly dusty. Landing sites were chosen in areas that appear to minimize dust using the TES Dust Cover Index.

Arcadia Planitia: Challenges

1st EZ Workshop for Human Missions to Mars



Acheron Fossae

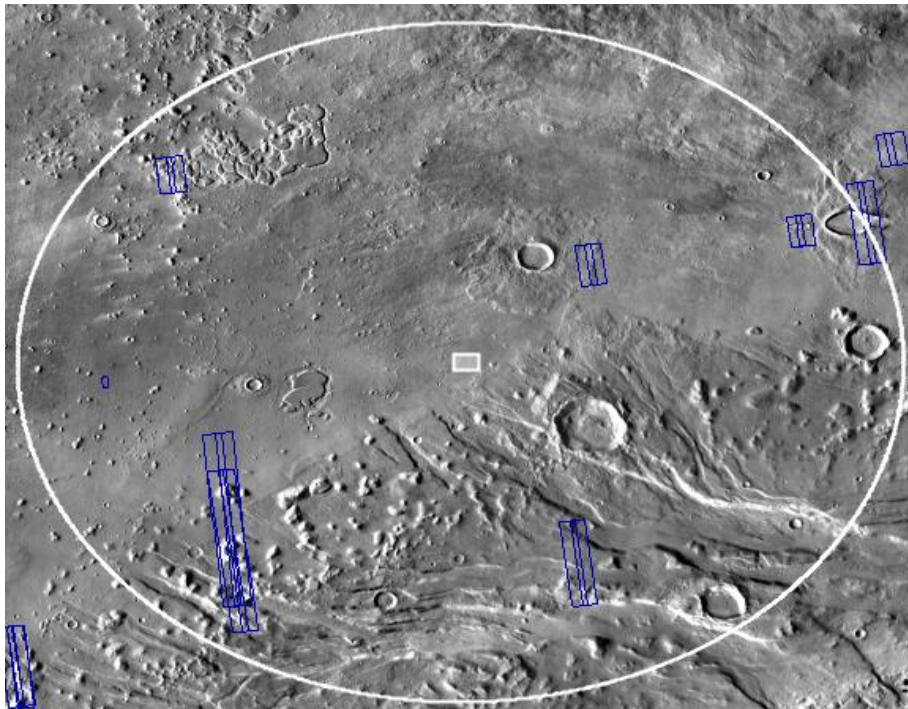


Erebus Montes

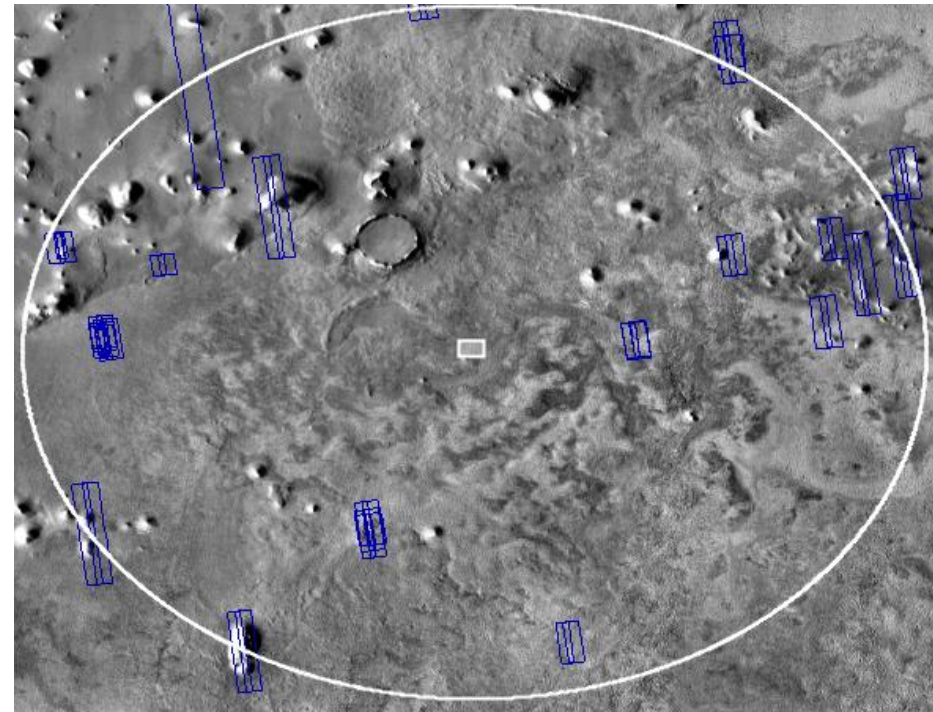
Composition: Very little CRISM data for either EZ.

Arcadia Planitia: Challenges

1st EZ Workshop for Human Missions to Mars



Acheron Fossae



Erebus Montes

High-resolution coverage: Some HiRISE coverage, but not much.



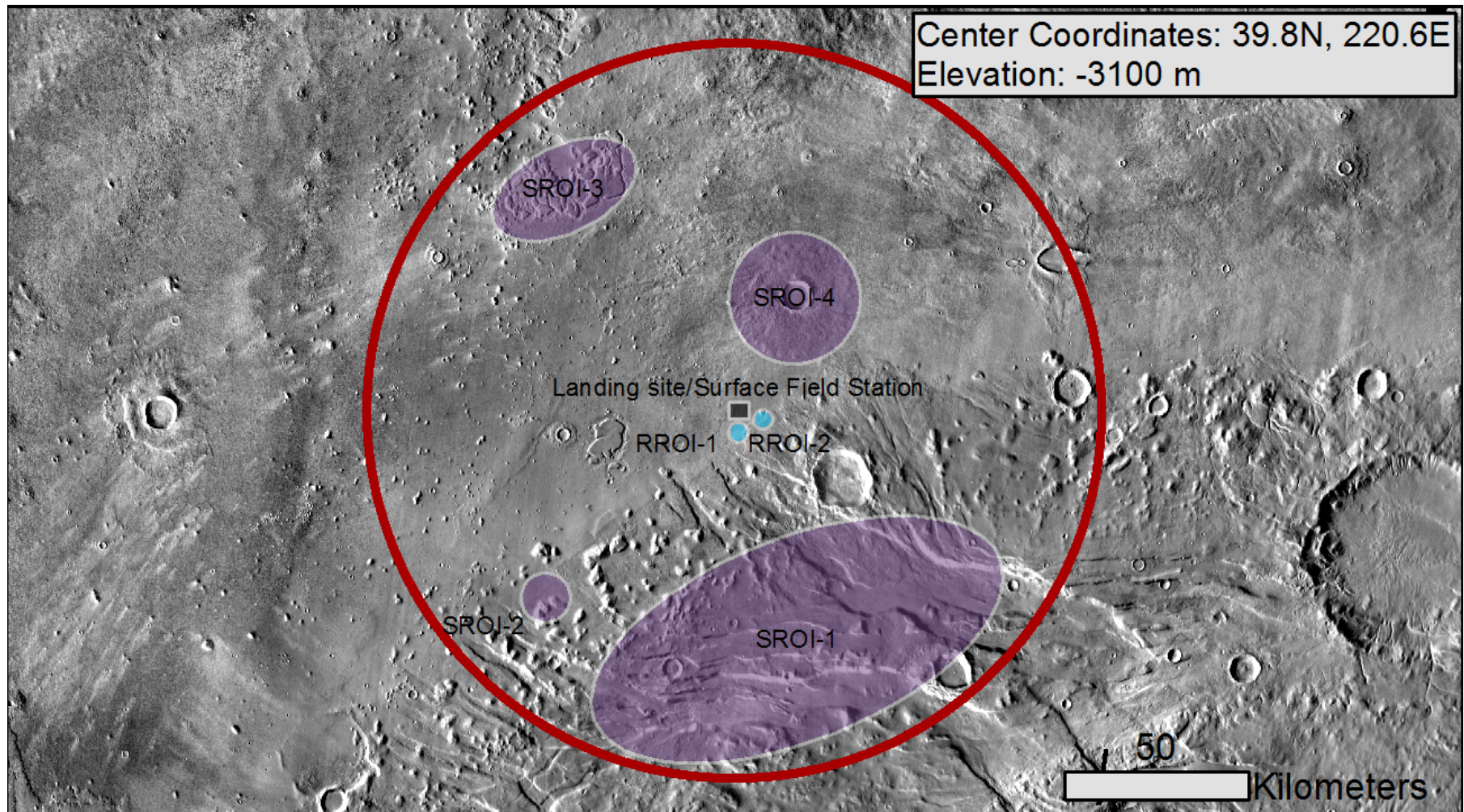
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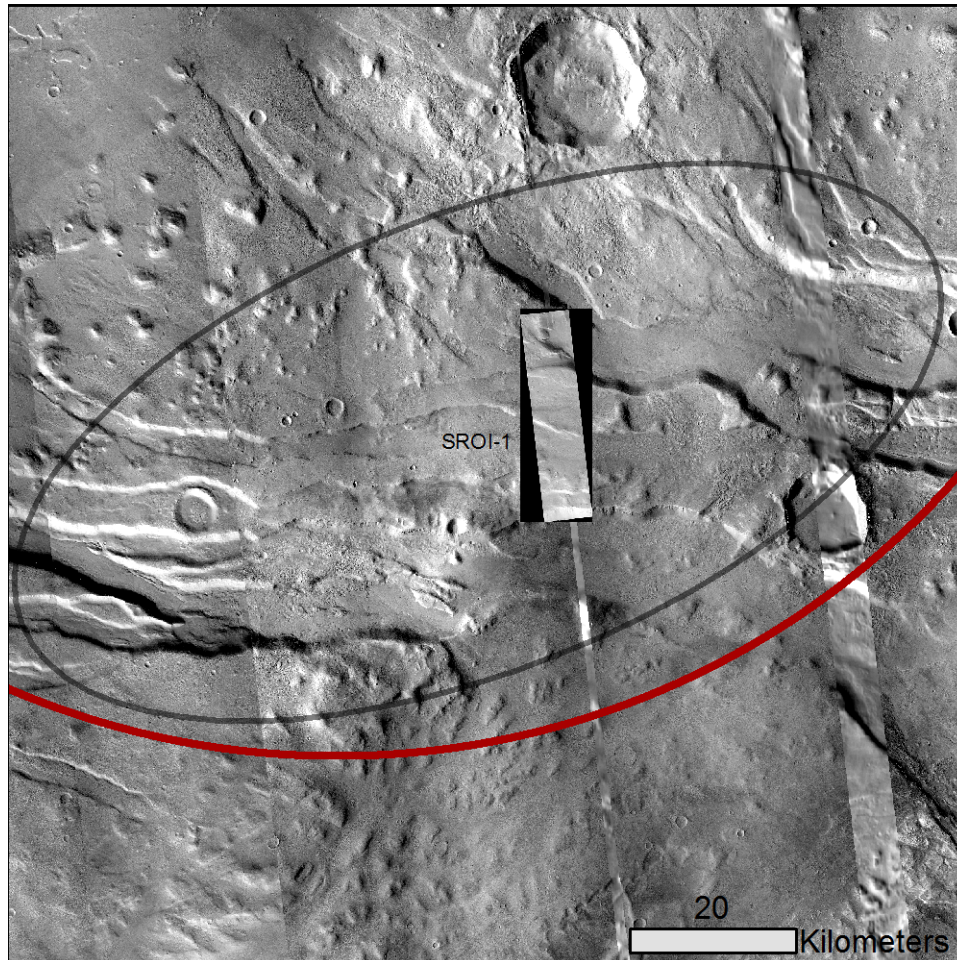
Exploration Zone Map

1st EZ Workshop for Human Missions to Mars



SCIENCE ROIs

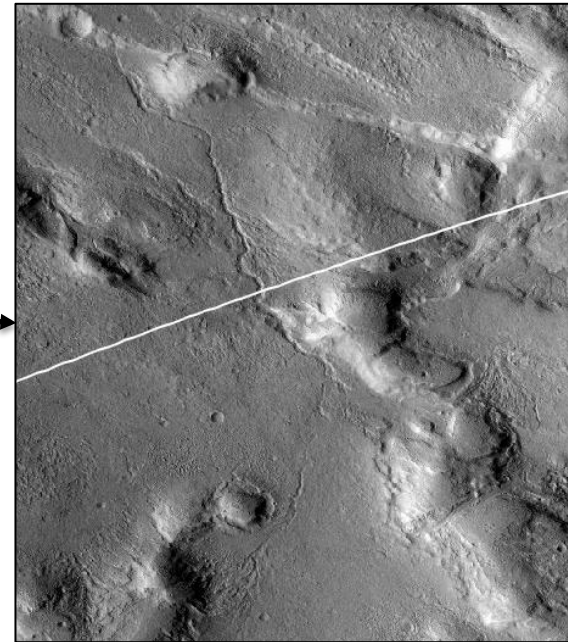
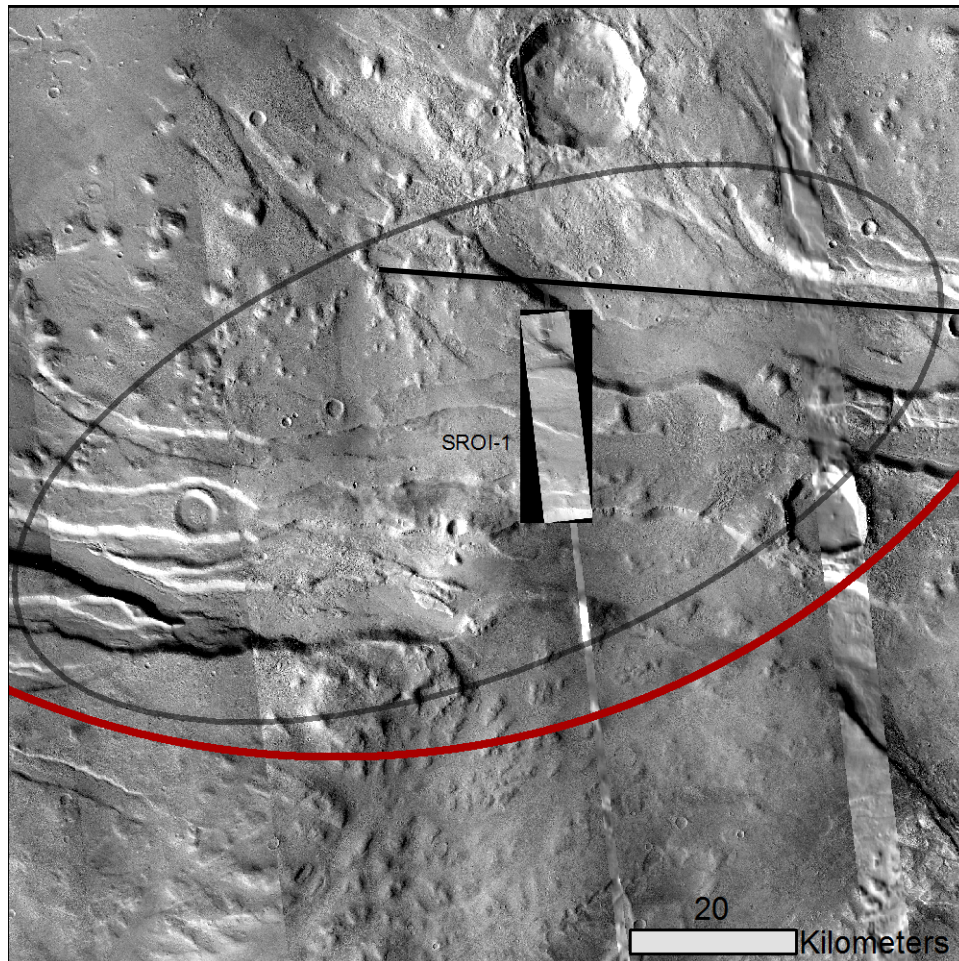
Science ROI 1



Acheron Fossae

- 38.7N, 221E. Well-covered in CTX, minimal HiRISE.
- Acheron Fossae graben system, late Noachian highlands terrain.
 - Presence of Noachian rocks.
 - Unit with regional/global extent.
 - Significant tectonic activity.
- Many different features of interest within this larger region.

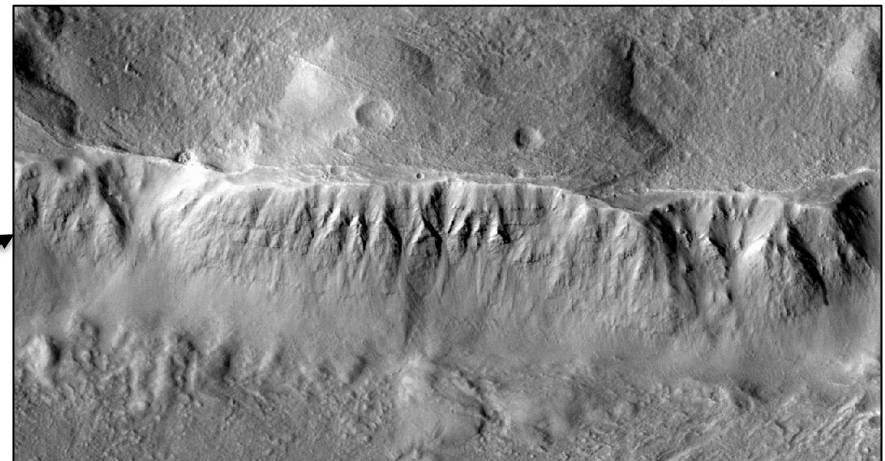
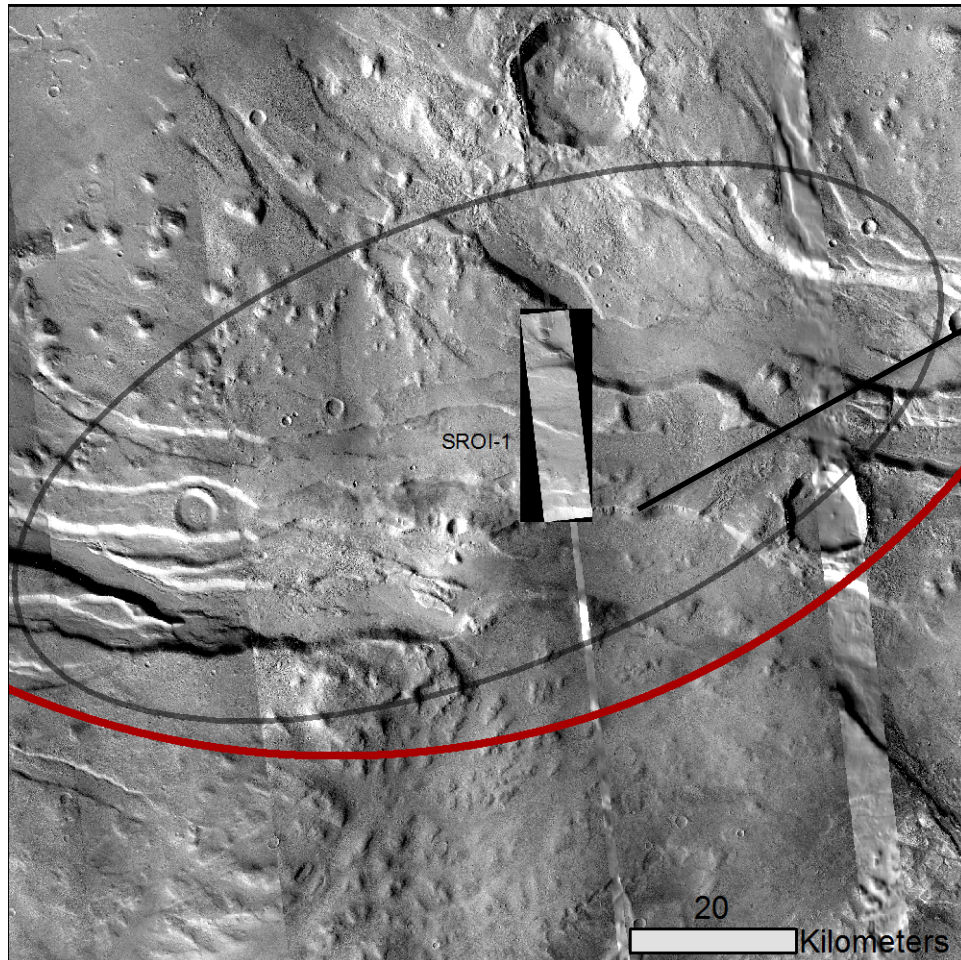
Science ROI 1



- Possible channel-like features in CTX:
 - Past aqueous activity?
 - Habitability potential?

Science ROI 1

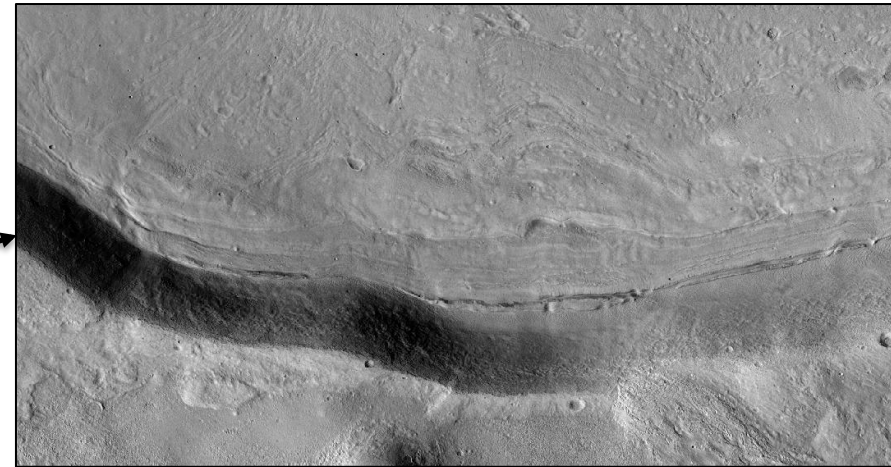
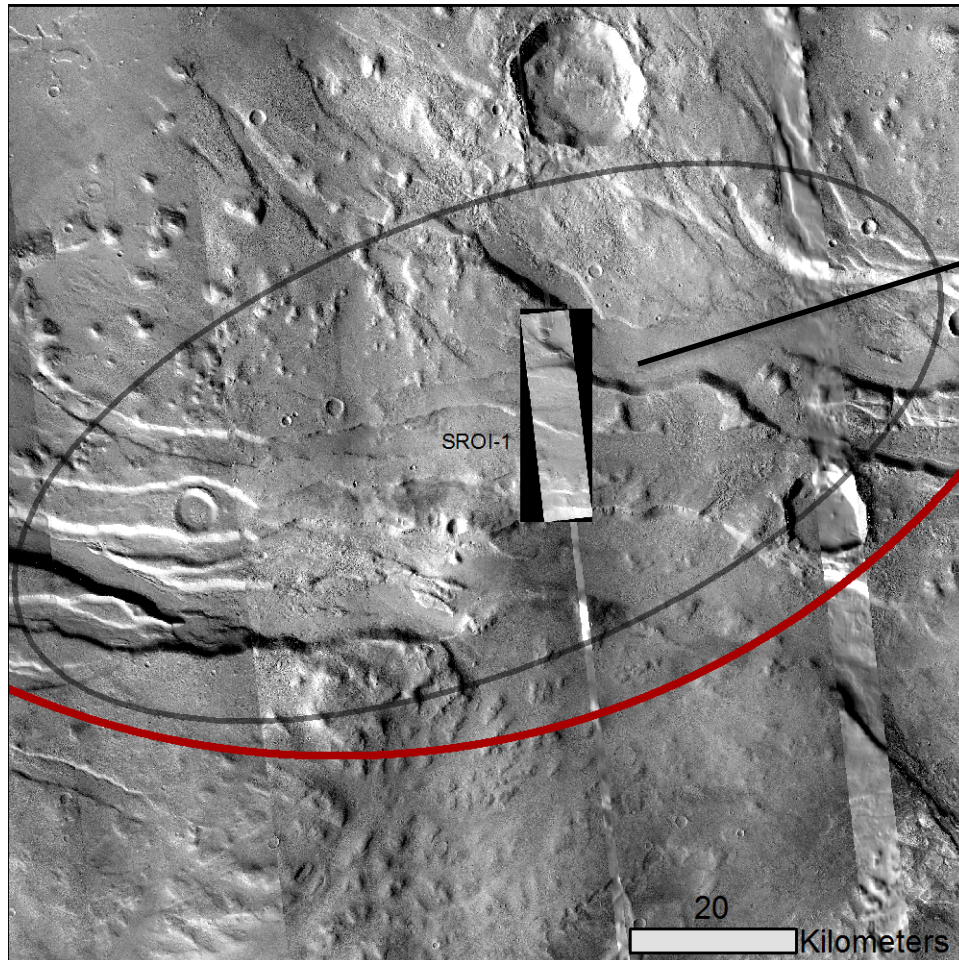
1st EZ Workshop for Human Missions to Mars



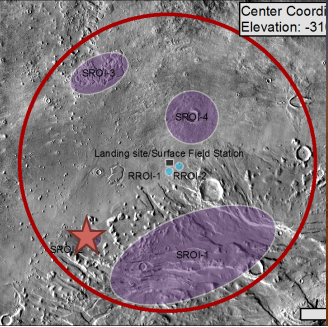
- Exposures along graben walls (multiple locations).
 - Access to Noachian bedrock units?

Science ROI 1

1st EZ Workshop for Human Missions to Mars

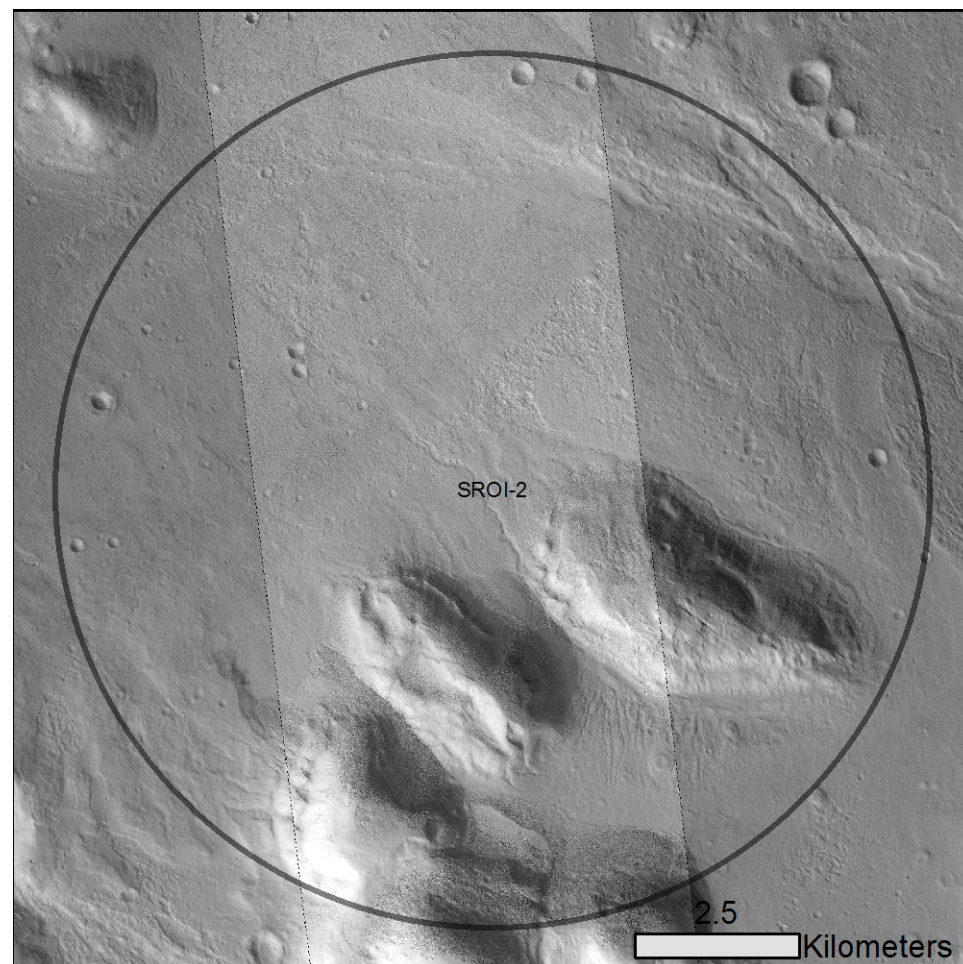


- Glacial fill on graben floors (likely Amazonian in age).
 - Range of surface ages (relative).
 - Access to Amazonian subsurface ice.

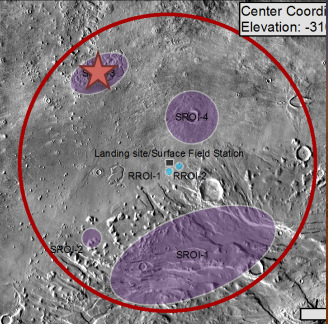


Science ROI 2

1st EZ Workshop for Human Missions to Mars

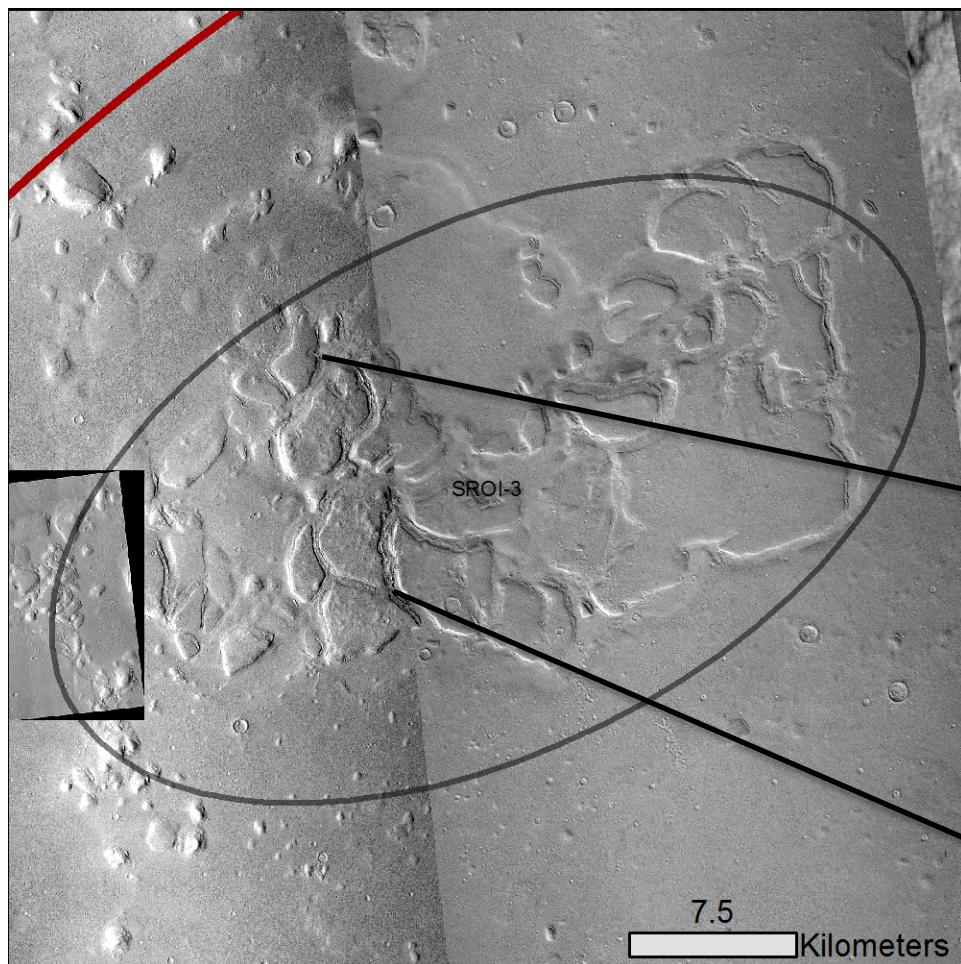


- 38.9N, 219.5E. Covered by CTX, HiRISE.
- Small channel-like feature, not in fluvial network – possible evidence of sub-glacial liquid water flow.
 - Aqueous activity
 - Potential habitability?

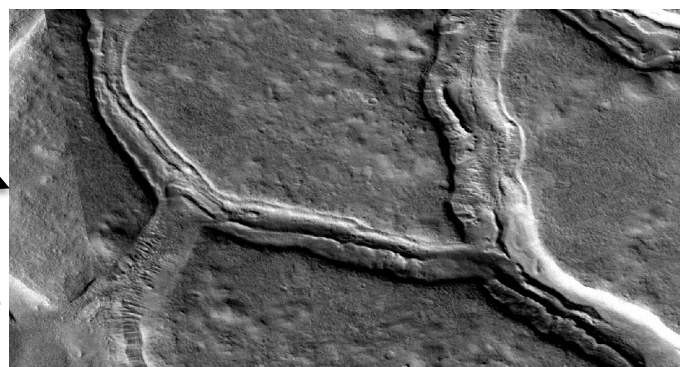


Science ROI 3

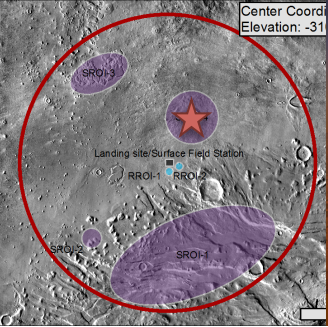
1st EZ Workshop for Human Missions to Mars



- 40.8N, 219.6E. CTX coverage, little HiRISE.
- Periglacial-type terrain; large blocky mesas
- Abundant subsurface ice (Amazonian), including glacial fill between blocks.

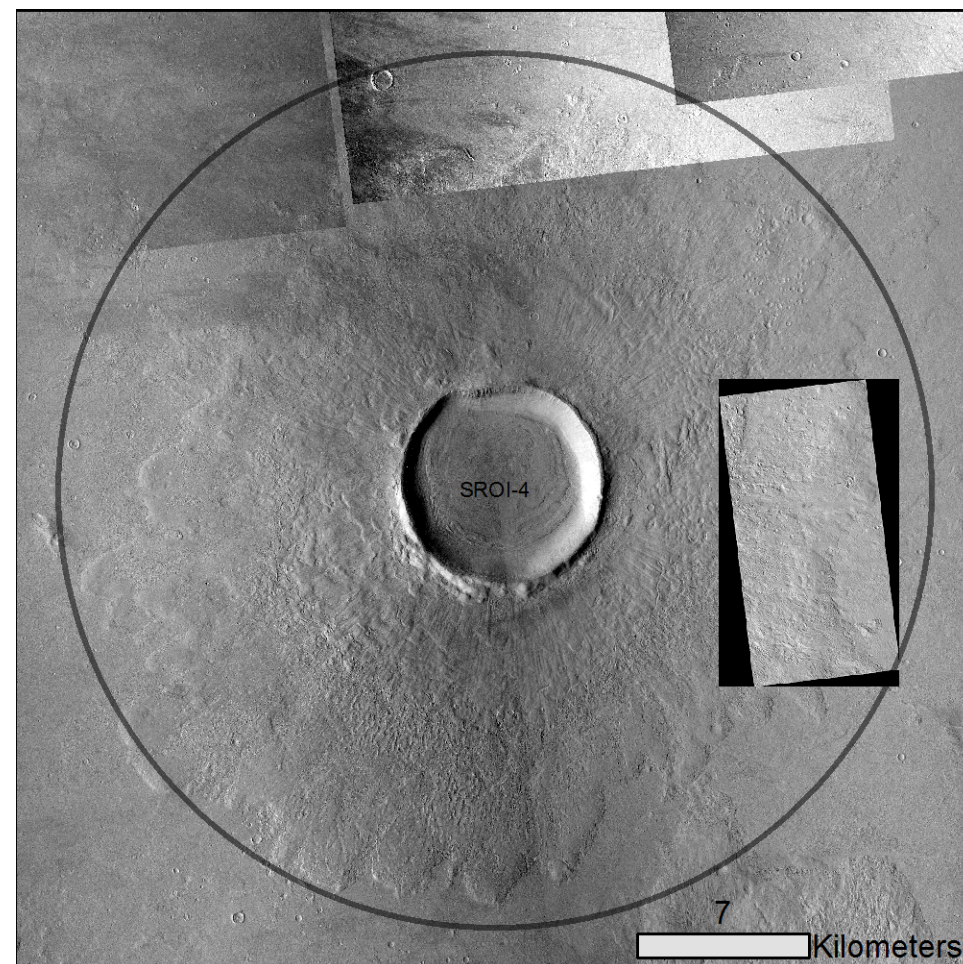


Acheron Fossae

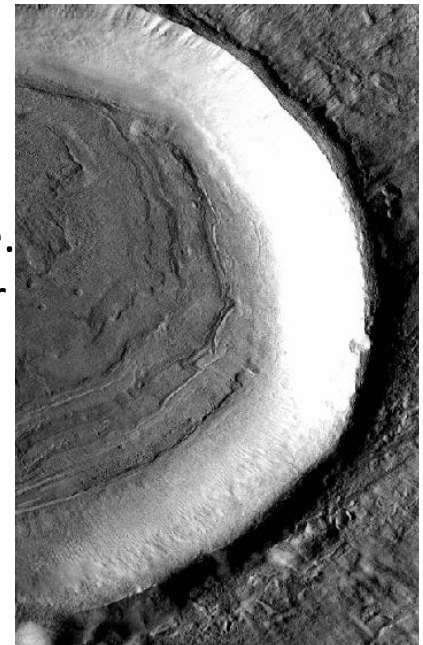


Science ROI 4

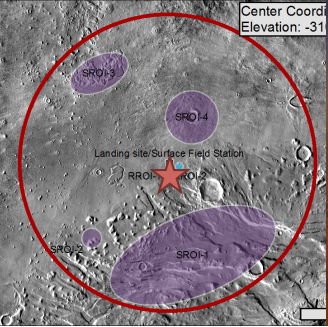
1st EZ Workshop for Human Missions to Mars



- 40.3N, 221E. CTX coverage, some HiRISE.
- Double layer ejecta impact crater, concentric crater fill, radial texture apparent in parts of ejecta.
 - Access to Amazonian-aged subsurface ice.
 - Primary crater deposits (with unique ejecta morphology).

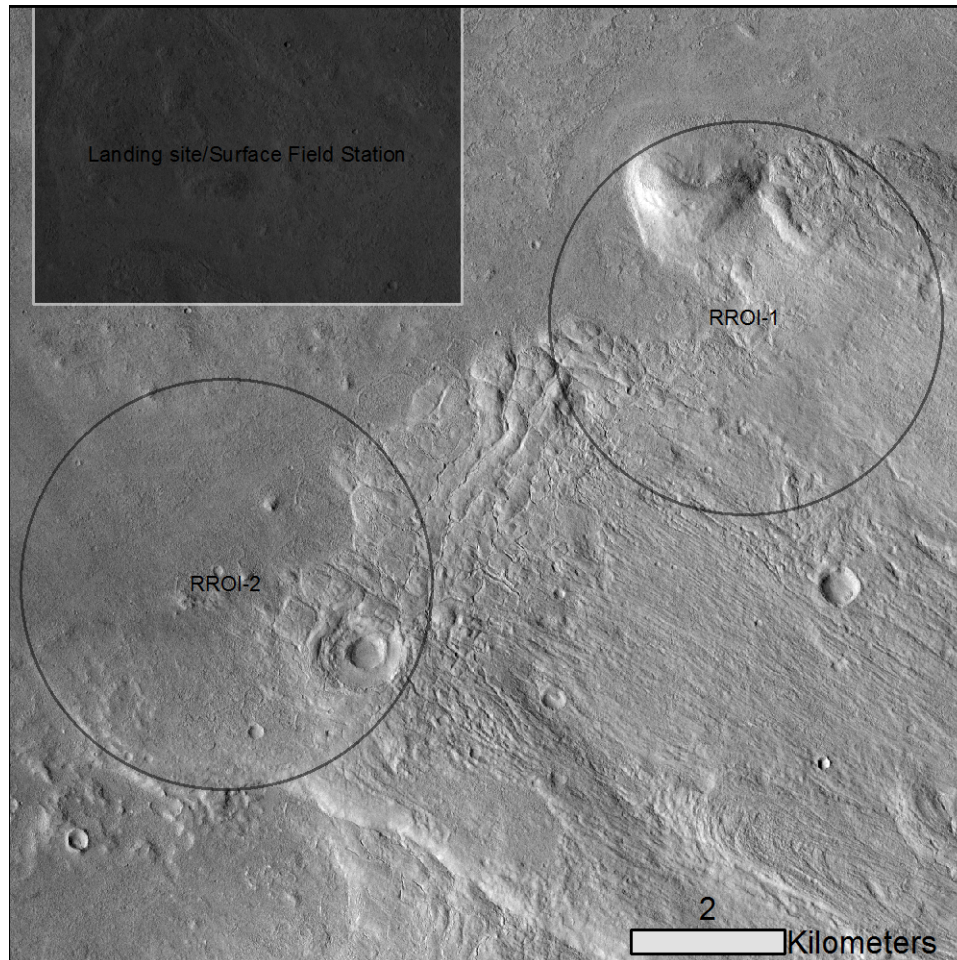


RESOURCE ROI_s



Resource ROIs

1st EZ Workshop for Human Missions to Mars



Acheron Fossae

- RROI-1: 39.8N, 220.8E
RROI-2: 39.7N, 220.6E
- Sample RROI locations near HZ, on the fringes of lobate, likely ice-rich, ejecta material.
- Excess water ice thought to be abundant throughout this entire region.

Science ROI(s) Rubric

1st EZ Workshop for Human Missions to Mars

Site Factors					SROI1	SROI2	SROI3	SROI4	RROI1	RROI2	EZ SUM	
Science Site Criteria	Astrobio	Threshold	AND/OR	Potential for past habitability	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	0,2	
				Potential for present habitability/refugia	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	0,6		
		Qualifying	Potential for organic matter, w/ surface exposure			<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	
	Atmospheric Science	Threshold	Noachian/Hesperian rocks w/ trapped atmospheric gases			<div></div>	<div></div>					0,2
			Meteorological diversity in space and time			<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	
		Qualifying	High likelihood of surface-atmosphere exchange			<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	6,0
			Amazonian subsurface or high-latitude ice or sediment			<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	6,0
			High likelihood of active trace gas sources			<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	
	Geoscience	Threshold	Range of martian geologic time; datable surfaces			<div></div>		<div></div>				2,0
			Evidence of aqueous processes			<div></div>	<div></div>					2,0
			Potential for interpreting relative ages			<div></div>	<div></div>	<div></div>	<div></div>			2,0
		Qualifying	Igneous Rocks tied to 1+ provinces or different times			<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	
			Near-surface ice, glacial or permafrost			<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	6,0
			Noachian or pre-Noachian bedrock units			<div></div>	<div></div>					2,0
			Outcrops with remnant magnetization			<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	
			Primary, secondary, and basin-forming impact deposits			<div></div>	<div></div>	<div></div>	<div></div>			4,0
			Structural features with regional or global context			<div></div>						1,0
	Diversity of aeolian sediments and/or landforms			<div></div>	<div></div>	<div></div>	<div></div>	<div></div>	<div></div>			


Key	
●	Yes
○	Partial Support or Debated
	No
?	Indeterminate

Resource ROI(s) Rubric

1st EZ Workshop for Human Missions to Mars

Site Factors				SROI1	SROI2	SROI3	SROI4	RROI1	RROI2	EZ SUM
ISRU and Civil Engineering Criteria	Engineering		Meets First Order Criteria (Latitude, Elevation, Thermal Inertia)		●	●	●	●	●	6,0
	Water Resource	Threshold	AND/OR	Potential for ice or ice/regolith mix	●	●	●	●	●	6,0
				Potential for hydrated minerals	?	?	?	?	?	
				Quantity for substantial production	●	●	●	●	●	6,0
				Potential to be minable by highly automated systems	●	●	●	●	●	6,0
				Located less than 3 km from processing equipment site				●	●	2,0
				Located no more than 3 meters below the surface	●	●	●	●	●	6,0
				Accessible by automated systems				●	●	2,0
		Qualifying		Potential for multiple sources of ice, ice/regolith mix and hydrated minerals						
				Distance to resource location can be >5 km						
				Route to resource location must be (plausibly) traversable						
	Civil Engineering	Threshold		~50 sq km region of flat and stable terrain with sparse rock distribution						
				1–10 km length scale: <10°	○	○	●	○	●	3,3
				Located within 5 km of landing site location				●	●	2,0
		Qualifying		Located in the northern hemisphere	●	●	●	●	●	6,0
				Evidence of abundant cobble sized or smaller rocks and bulk, loose regolith	?	?	?	?	?	
	Food Production	Qualifying		Utilitarian terrain features	?	?	?	?	?	
				Low latitude						
				No local terrain feature(s) that could shadow light collection facilities						
				Access to water	●	●	●	●	●	6,0
	Metal/Silicon Resource	Threshold		Access to dark, minimally altered basaltic sands	?	?	?	?	?	
				Potential for metal/silicon	?	?	?	?	●	2,0
				Potential to be minable by highly automated systems				?	?	
				Located less than 3 km from processing equipment site				●	●	2,0
				Located no more than 3 meters below the surface				○	○	0,2
				Accessible by automated systems				?	?	
		Qualifying		Potential for multiple sources of metals/silicon						
				Distance to resource location can be >5 km						
				Route to resource location must be (plausibly) traversable						

Key	
●	Yes
○	Partial Support or Debated
	No
?	Indeterminate



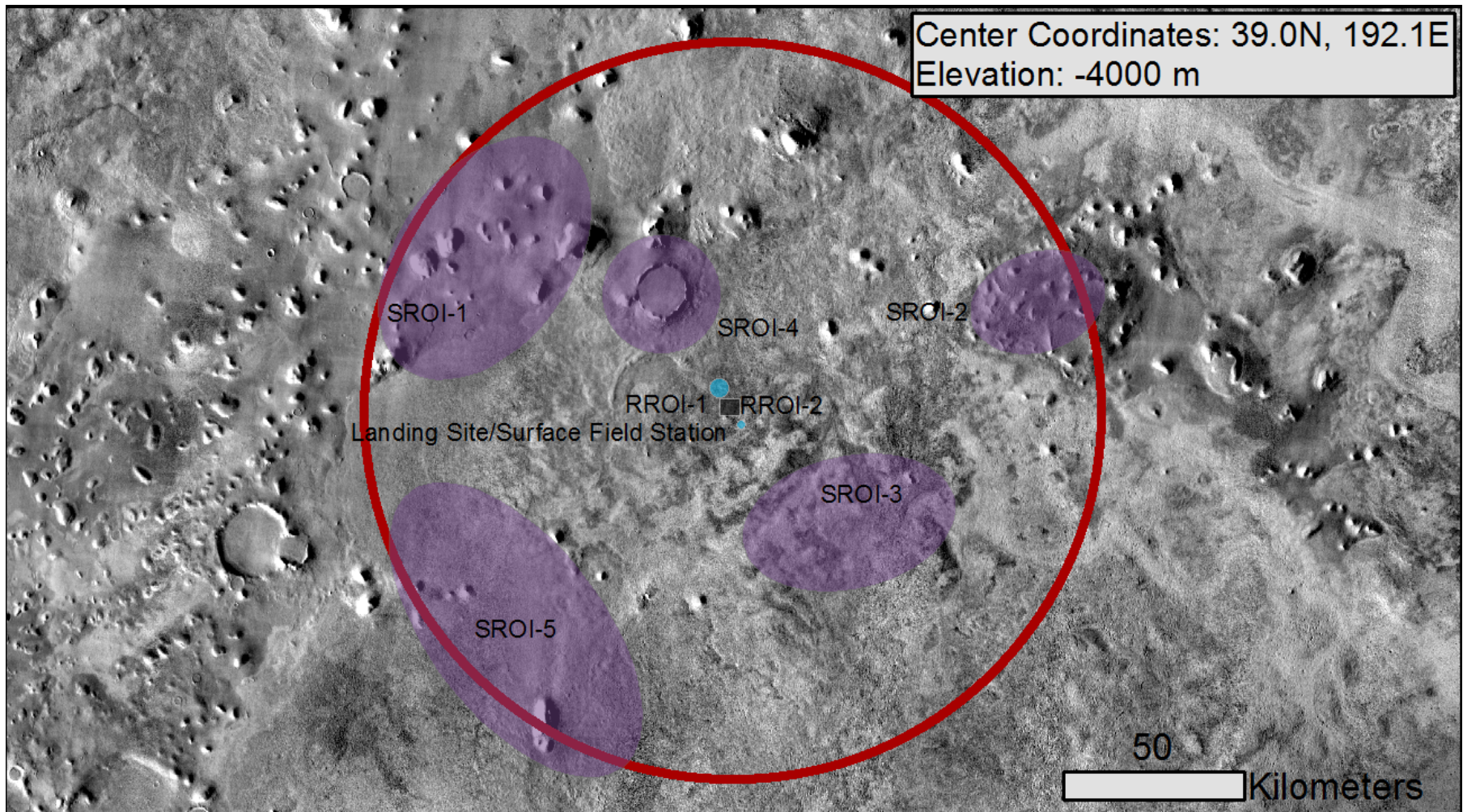
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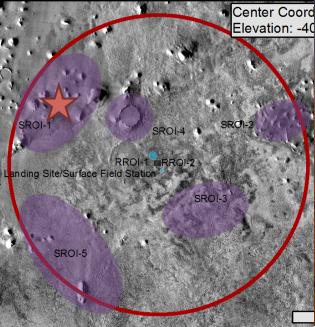
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Exploration Zone Map

1st EZ Workshop for Human Missions to Mars

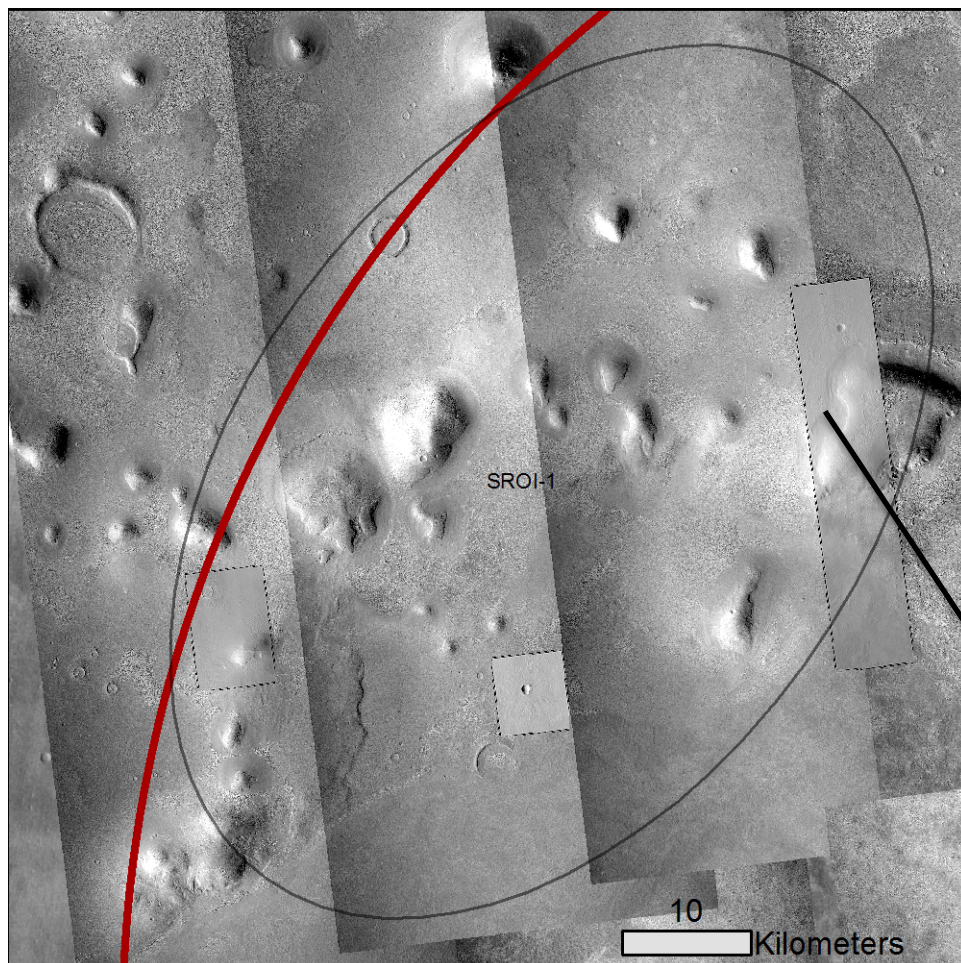


SCIENCE ROIs



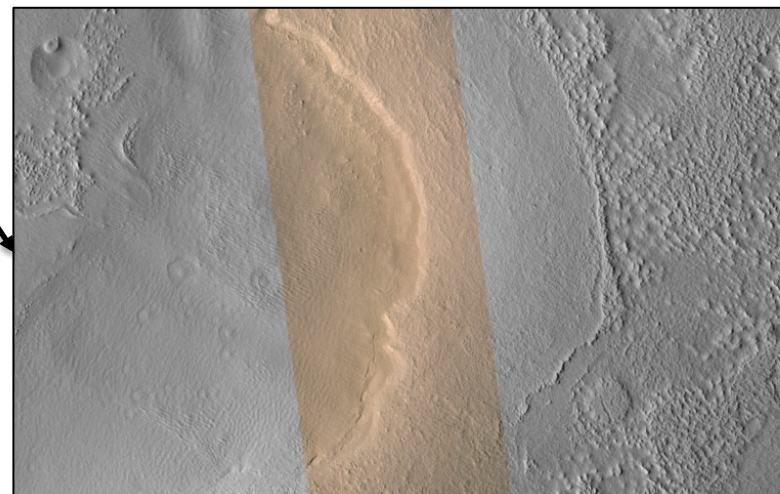
Science ROI 1

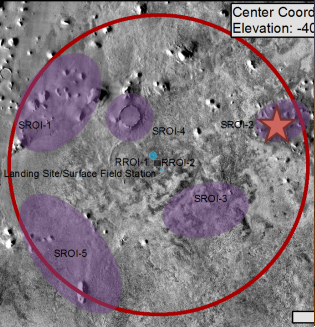
1st EZ Workshop for Human Missions to Mars



Erebus Montes

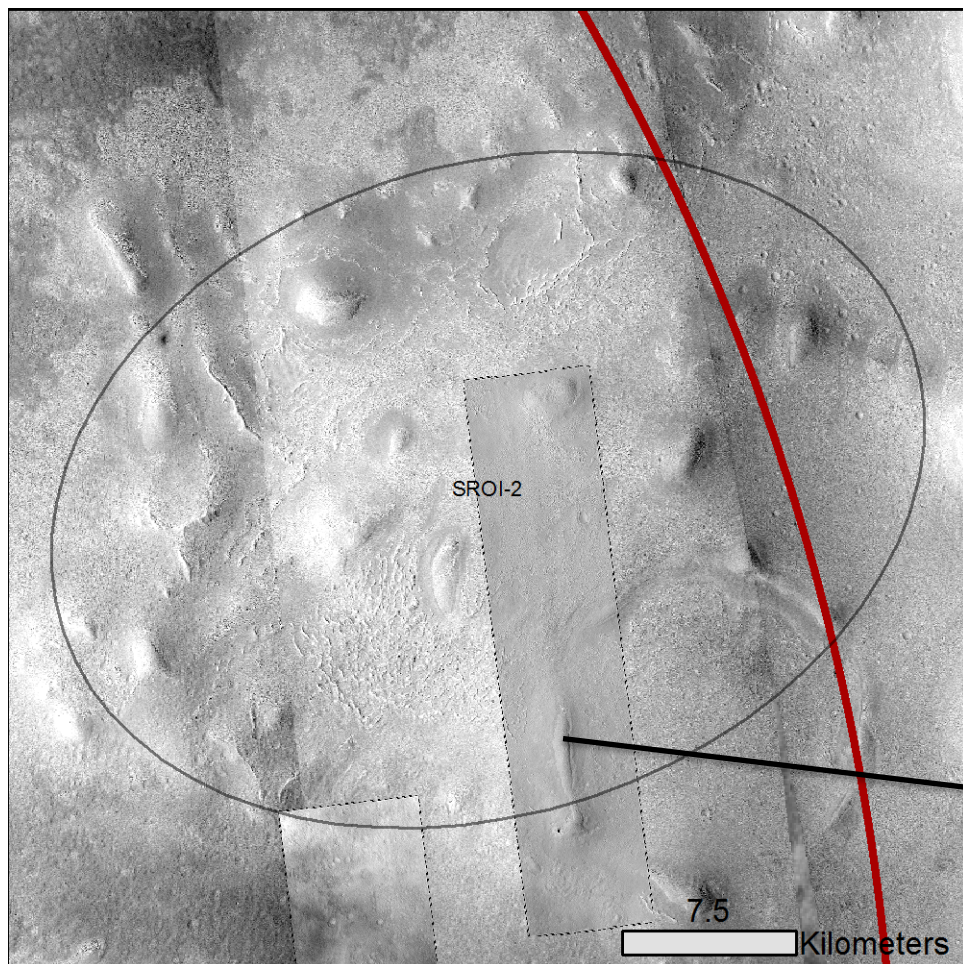
- 39.7N, 190.6E. Well-covered in CTX, some HiRISE.
- Exposures of Hesperian-Noachian transition terrain.
 - Noachian/Hesperian rocks?
- Some mounds of HNT material have ice-rich lobate debris aprons.
 - Amazonian-aged subsurface ice.
 - Relative ages.
- Located near a recent ice-exposing impact crater.
 - Subsurface ice.
 - Impact crater processes.





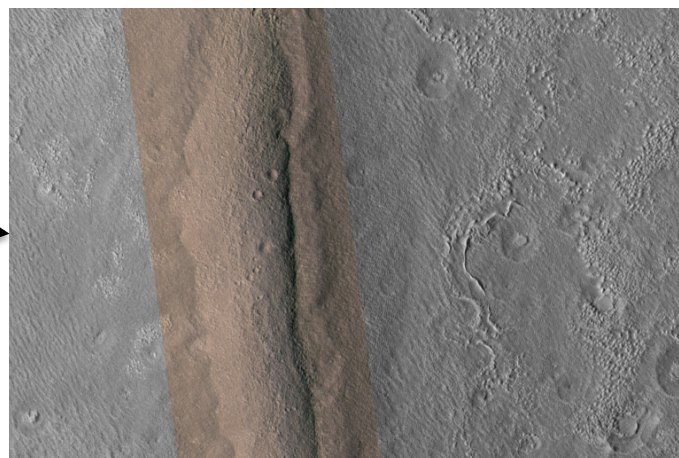
Science ROI 2

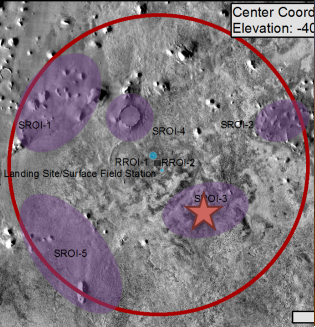
1st EZ Workshop for Human Missions to Mars



Erebus Montes

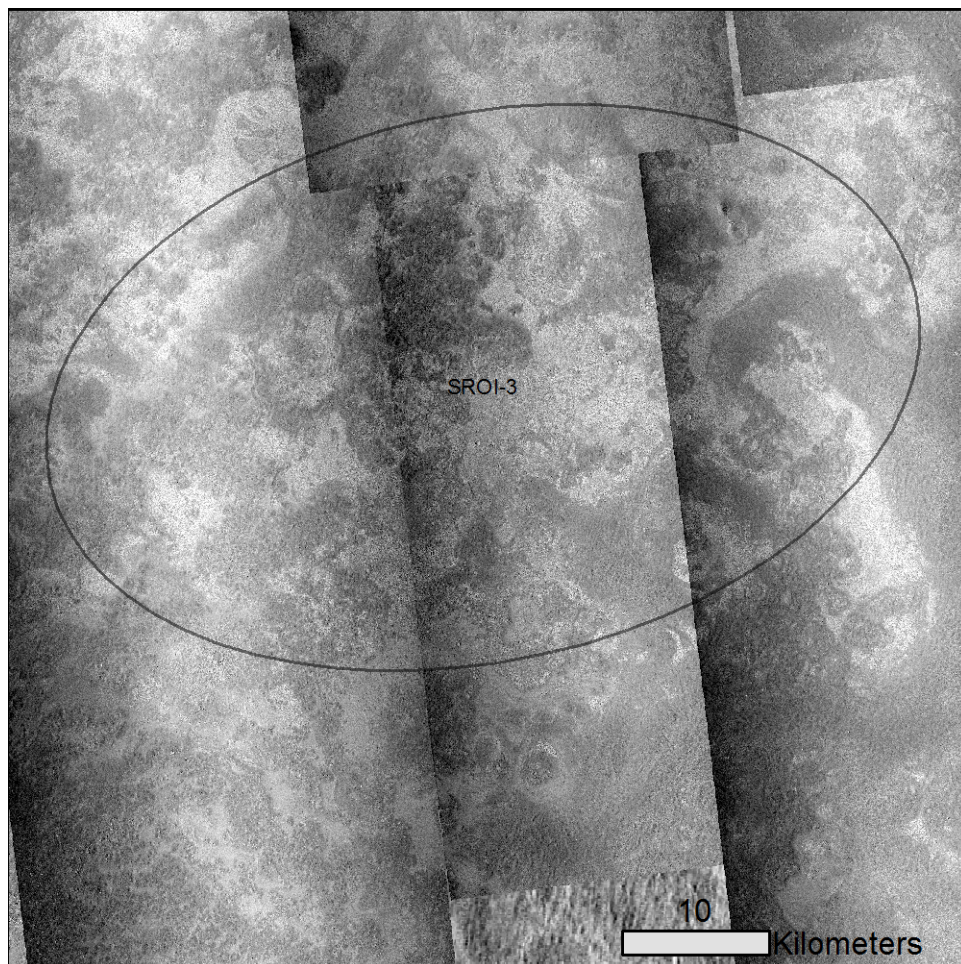
- 39.5N, 194E. Well-covered in CTX, some HiRISE.
- Additional exposures of Hesperian-Noachian transition terrain.
- Some mounds of HNT material have ice-rich lobate debris aprons. Filled/buried crater seen in SE region.
 - Impact processes, relative ages.





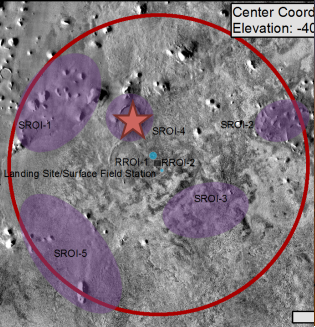
Science ROI 3

1st EZ Workshop for Human Missions to Mars



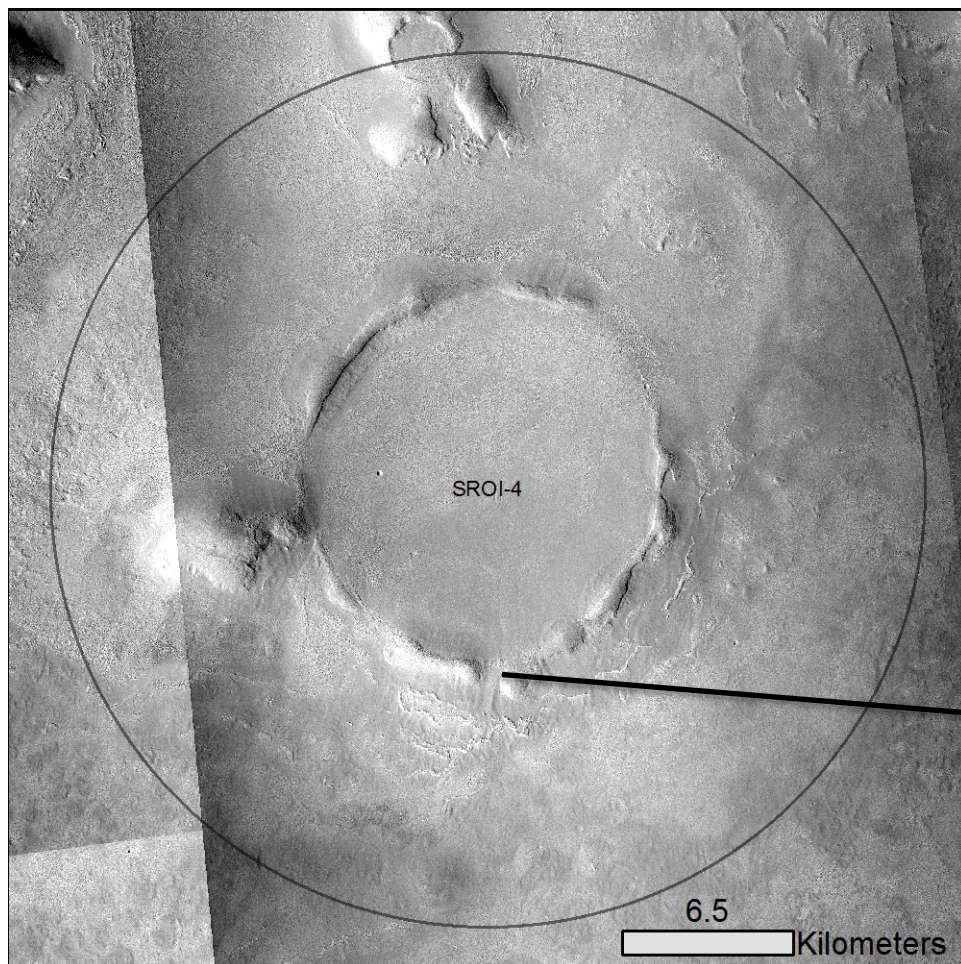
Erebus Montes

- 38.5N, 192.8E. Well-covered in CTX, no HiRISE.
- Contact between multiple episodes of Amazonian volcanism.
 - Relative surface ages.
 - Possibility for lava tube caves?
- Also includes some evidence for subsurface ice.



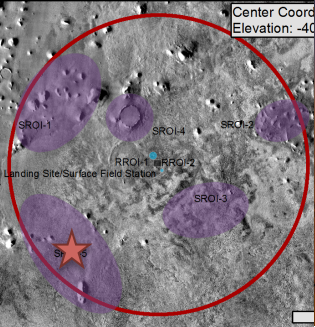
Science ROI 4

1st EZ Workshop for Human Missions to Mars



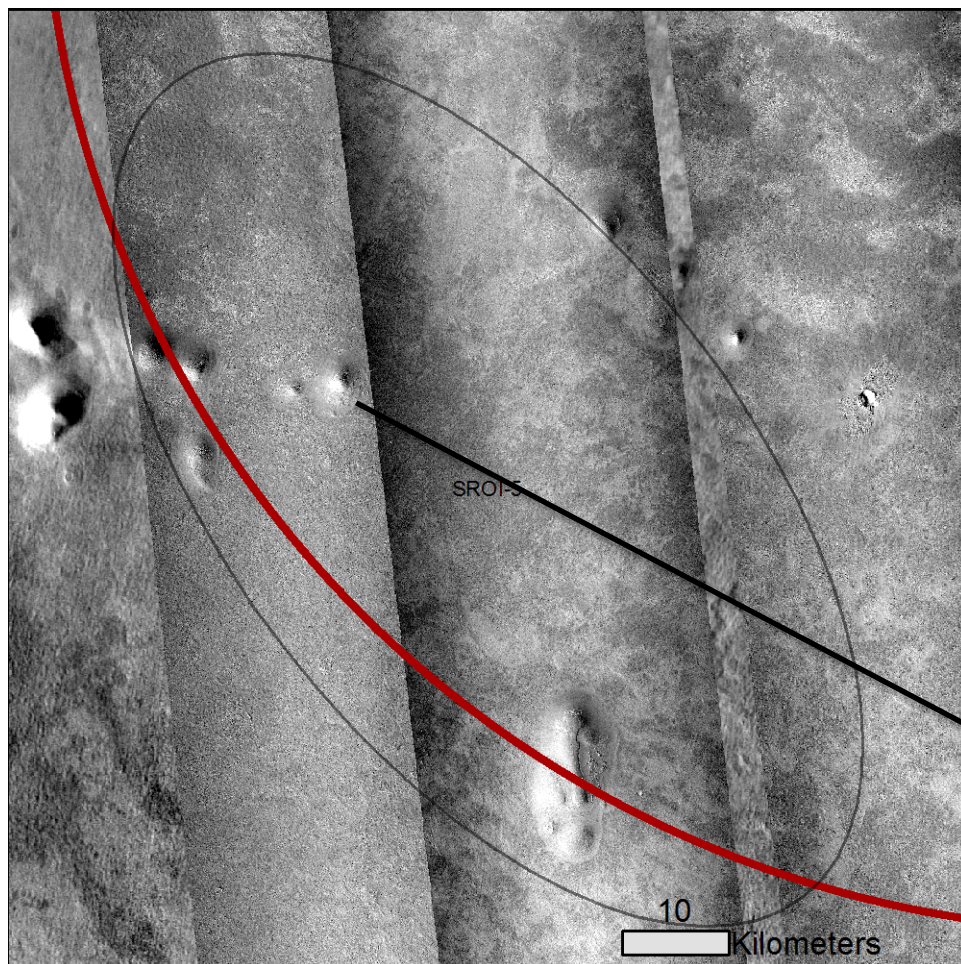
- 39.5N, 191.5E. Well-covered in CTX, no HiRISE.
- Filled/buried impact crater, evidence for glacial/periglacial processes.



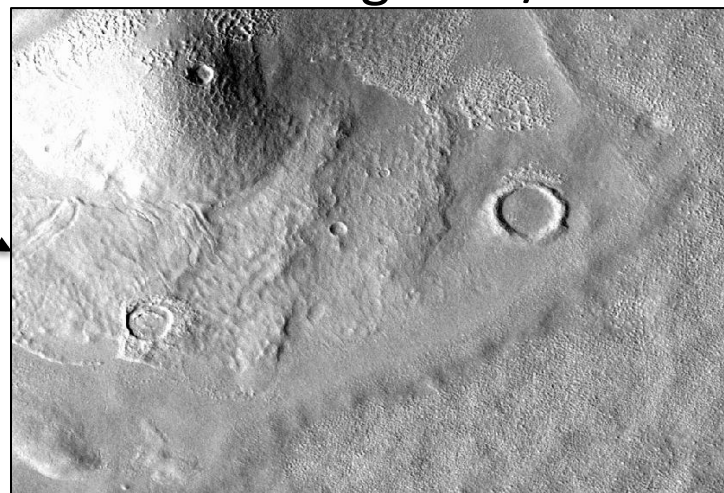


Science ROI 5

1st EZ Workshop for Human Missions to Mars

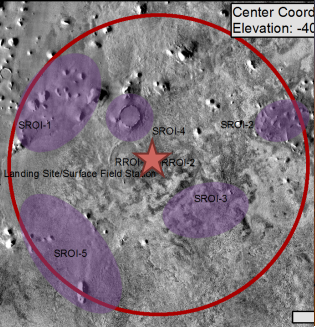


- 37.9N, 190.8E. Well-covered in CTX, some HiRISE.
- Mounds with lobate debris aprons, surrounded by Amazonian lava flows. Evidence for glacial/



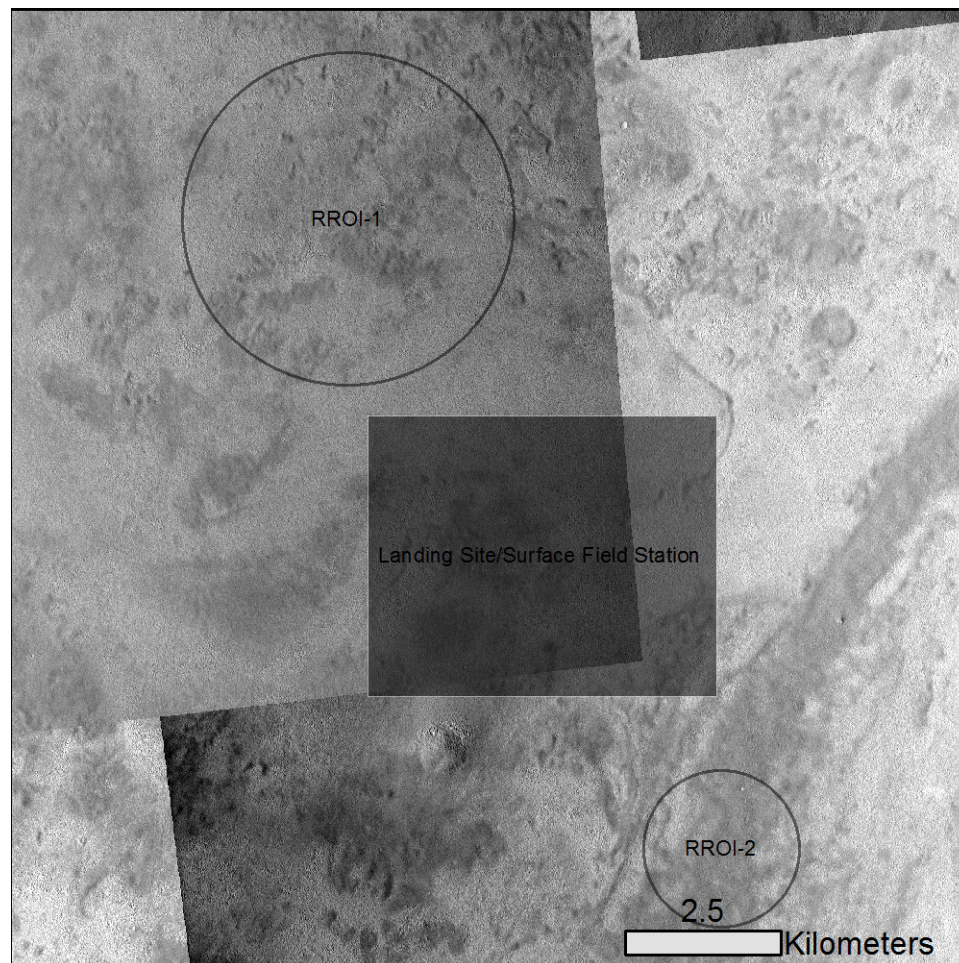
Erebus Montes

RESOURCE ROI_s



Resource ROIs

1st EZ Workshop for Human Missions to Mars



Erebus Montes

- RROI-1: 39.1N, 192E
RROI-2: 38.9N, 192.2E
- Sample RROI locations near HZ. Both contain evidence for excess subsurface water ice (thought to be abundant throughout this entire region).

Science ROI(s) Rubric

1st EZ Workshop for Human Missions to Mars

Site Factors					SROI1	SROI2	SROI3	SROI4	SROI5	RROI1	RROI2	EZ SUM
Science Site Criteria	Astrobio	Threshold	AND/OR	Potential for past habitability	?	?	?	?	?	?	?	
				Potential for present habitability/refugia	○	○	○	○	○	○	0,7	
		Qualifying		Potential for organic matter, w/ surface exposure	?	?	?	?	?	?	?	
	Atmospheric Science	Threshold		Noachian/Hesperian rocks w/ trapped atmospheric gases	○	○			○			0,3
		Qualifying		Meteorological diversity in space and time	?	?	?	?	?	?	?	
				High likelihood of surface-atmosphere exchange	●	●	●	●	●	●	●	7,0
				Amazonian subsurface or high-latitude ice or sediment	●	●	●	●	●	●	●	7,0
				High likelihood of active trace gas sources								
	Geoscience	Threshold		Range of martian geologic time; datable surfaces	●	●	●	●	●			5,0
				Evidence of aqueous processes								
				Potential for interpreting relative ages	●	●	●	●	●			5,0
		Qualifying		Igneous Rocks tied to 1+ provinces or different times	?	?	?	?	?	?	?	
				Near-surface ice, glacial or permafrost	●	●	●	●	●	●	●	7,0
				Noachian or pre-Noachian bedrock units	○	○			○			0,3
				Outcrops with remnant magnetization	?	?	?	?	?	?	?	
				Primary, secondary, and basin-forming impact deposits	●	●	●	●	●			5,0
				Structural features with regional or global context	●	●			●			3,0
			Diversity of aeolian sediments and/or landforms	?	?	?	?	?	?	?		

Key	
●	Yes
○	Partial Support or Debated
	No
?	Indeterminate

Resource ROI(s) Rubric

1st EZ Workshop for Human Missions to Mars

Site Factors					SROI1	SROI2	SROI3	SROI4	SROI5	RROI1	RROI2	EZ SUM	
ISRU and Civil Engineering Criteria	Engineering		Meets First Order Criteria (Latitude, Elevation, Thermal Inertia)		●	●	●	●	●	●	●	7,0	
	Water Resource	Threshold	AND/ OR	Potential for ice or ice/regolith mix	●	●	●	●	●	●	●	7,0	
				Potential for hydrated minerals	?	?	?	?	?	?	?		
			Quantity for substantial production	●	●	●	●	●	●	●	7,0		
			Potential to be minable by highly automated systems	●	●	●	●	●	●	●	7,0		
			Located less than 3 km from processing equipment site						●	●	2,0		
			Located no more than 3 meters below the surface	●	●	●	●	●	●	●	7,0		
			Accessible by automated systems						●	●	2,0		
		Qualifying	Potential for multiple sources of ice, ice/regolith mix and hydrated minerals										
			Distance to resource location can be >5 km										
			Route to resource location must be (plausibly) traversable										
	Civil Engineering	Threshold	~50 sq km region of flat and stable terrain with sparse rock distribution										
			1–10 km length scale: <10°		○	●	●	○	●	●	●	5,2	
			Located within 5 km of landing site location							●	●	2,0	
		Qualifying	Located in the northern hemisphere		●	●	●	●	●	●	●	7,0	
			Evidence of abundant cobble sized or smaller rocks and bulk, loose regolith		?	?	?	?	?	?	?		
	Food Production	Qualifying	Utilitarian terrain features		?	?	?	?	?	?	?	?	
			Low latitude										
			No local terrain feature(s) that could shadow light collection facilities										
			Access to water		●	●	●	●	●	●	●	●	7,0
	Metal/Silicon Resource	Threshold	Access to dark, minimally altered basaltic sands		?	?	?	?	?	?	?	?	
			Potential for metal/silicon		?	?	?	?	?	●	●	2,0	
			Potential to be minable by highly automated systems							●	●	2,0	
			Located less than 3 km from processing equipment site							●	●	2,0	
			Located no more than 3 meters below the surface							○	○	0,2	
		Accessible by automated systems							?	?			
		Qualifying	Potential for multiple sources of metals/silicon										
			Distance to resource location can be >5 km										
			Route to resource location must be (plausibly) traversable										

Key	
●	Yes
○	Partial Support or Debated
	No
?	Indeterminate

BONUS Science ROI

1st EZ Workshop for Human Missions to Mars



For both Acheron Fossae and Erebus Montes: The Arcadia Planitia Ice Sheet

- Easy access to subsurface ice means that it should be easy to sample and study as well.
- Science questions:
 - How did the ice sheet form?
 - How long did the ice deposition take, and how long has it existed?
 - How thick is the ice sheet?
 - How is the ice interacting with the atmosphere?
 - Did the ice preserve organics or other biosignatures?
 - Could the ice have contained habitable microenvironments now or in the past?

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Water, water (ice) everywhere – but not all of it's to drink!

Highest Priority EZ Data Needs

1st EZ Workshop for Human Missions to Mars



- Science Priority: surface composition data
 - Surface dust affects CRISM observations.
 - Strategy: Target CRISM observations in regions with slightly less dust.
- Resource Priority: higher resolution of near-surface ice deposits, including spatial distribution
 - SAR RADAR, higher resolution neutron spectrometer